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LaBr<sub>3</sub>:Ce and CeBr<sub>3</sub> scintillators are candidates for gamma ray diagnostics at JET during deuterium-tritium campaigns when high count rates are expected. Both crystals are characterised by short scintillation decay times ~15-25 ns and a similar energy resolution ~3% at 1 MeV. The naturally occurring <sup>138</sup>La radioisotope (0.09%) present in LaBr<sub>3</sub>:Ce crystals explains the higher intrinsic background in comparison with CeBr<sub>3</sub>.

We report on a comparison of the internal activity of  $LaBr_3$ :Ce and  $CeBr_3$  scintillators of the same size: 3"×3". The scintillators were coupled to a Hamamatsu R6233 photomultiplier (PMT). Signals from the PMT were sent to a spectroscopy amplifier ORTEC 672 and finally to a TUKAN 8K USB multichannel analyser.

In Fig. 1 measured spectra normalised to the same live time are presented. The spectra are divided into two regions: a low energy interval up to 2.5 MeV (upper part) and a high energy region from 3 MeV (lower part).



Fig. 1. Comparison of natural background and internal activity for LaBr<sub>3</sub> and CeBr<sub>3</sub>. In the upper part the low energy spectrum up to 2.5 MeV is shown, in the lower part the high energy region from 3 MeV is presented.

At the energy range above ~4 MeV (lower figure) no structure is observed. This is important because in deuterium-tritium experiments observation of a 4.4 MeV gamma ray gives evidence that the reaction of interest has occurred.

$$^{9}Be + \alpha \rightarrow {}^{13}C \xrightarrow{n}{}^{12}C \stackrel{*}{\longrightarrow} \stackrel{\gamma(4.44 \, MeV)}{\longrightarrow} {}^{12}C \, as$$

The biggest differences in spectra registered with the LaBr<sub>3</sub>:Ce and CeBr<sub>3</sub> scintillators are observed for lower energies, see Fig. 2 for details.



Fig. 2. Response of  $3'' \times 3''$  CeBr<sub>3</sub> and LaBr<sub>3</sub>:Ce scintillators to natural and intrinsic background. Events due to internal contamination by actinides, observed between 1.5 MeV and 2.5 MeV, are seen for both scintillators. In the case of LaBr<sub>3</sub>:Ce, events connected with the presence of <sup>138</sup>La are registered for energies below 1.5 MeV.

 $^{138}$ La decays by an electron capture (EC) or  $\beta^{\circ}$ . EC is followed by emission of a 1.436 MeV  $\gamma$ -ray and X-ray cascade, giving a contribution to a peak at about 1.470 MeV. In addition,  $\beta^{\circ}$  decay produces a continuum above 0.789 MeV. An additional background in the energy region between 1.5 and 2.75 MeV is connected with a progeny of  $^{227}$ Ac contaminating both LaBr<sub>3</sub>:Ce and CeBr<sub>3</sub> scintillators.

Peaks originating from gamma transitions observed in a natural background (1.461 MeV from  $^{40}\text{K}$  and 2.615 MeV from  $^{208}\text{Tl}$ ) are also clearly seen. The 1.461 MeV gamma line overlaps with the 1.470 MeV gamma line from  $^{138}\text{La}$ .

In conclusion, CeBr<sub>3</sub> has the advantage in comparison with LaBr<sub>3</sub>:Ce because of its lower internal radioactivity below 1.5 MeV due to the lack of any radioactive isotopes.

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