



## **1. Gamma-ray Camera at JET**

The Gamma-ray Camera (GC) is one of the diagnostics system used at JET. It consists of 19 scintillator detectors, dedicated electronic devices and a data acquisition system. In each aluminum capsule a scintillator connected to a photomultiplier is fixed. Before the upgrade of GC, CsI scintillators and PIN-diode photomultipliers were used. These detectors are replaced by  $LaBr<sub>3</sub> scintillator$  and the MPPC device which is used as a photomultiplier due to project requirements. In the horizontal and vertical parts of GC 10 and 9 detectors are placed, respectively. The schematic of the GC (left) and a capsule (right) is shown in Fig. 1.



Fig. 1. Schematic of the JET Gamma-ray Camera (left) and a capsule with  $20\times15$  mm CeBr<sub>3</sub>based detector (right).





During the test phase in May 2015 a scintillator CeBr<sub>3</sub> with a MPPC array was placed in the channel #10 which is a part of the Horizontal Camera (HC). Data registered with this detector are analyzed in this work.

The electronic system consists of the FilterBox@NCBJ and MTCD@NCBJ devices. The MTCD@NCBJ device provides the power and the temperature compensation. The FilterBox@NCBJ is used for communication between the MTCD@NCBJ and the temperature sensors placed on the detectors.

The new Data Acquisition System (DAQ) was implemented and installed by the IST group.

### **2. Energy calibration**

The energy calibration measurements were performed on  $20<sup>th</sup>$  July, 2016. There was a series of one-hour measurements with a  $^{22}$ Na source, emitting 511 and 1274 keV. The energy calibration is obtained from a formula:

Energy (keV) = 
$$
(3.34 \cdot \text{channel} + 137)
$$
.

# **3. Analysis of raw data for runaway electron (RE) production from JET Gamma-ray Camera**

During analysis we extract time and energy for each event during a shot:

- read time and amplitude data from a DAQ file,
- perform energy calibration,
- use a proper code depending on a scintillator type  $(CsI)$  or  $CeBr<sub>3</sub>/LaBr<sub>3</sub>:Ce$  to obtain an amplitude and to find single (not pile-up) and pile-up events,
- apply a code AU\_main to unfold pile-up events,
- split events into two parts: with energy  $\leq 520 \text{ keV}$  and  $> 520 \text{ keV}$ . With such a value, all events from a 511 keV peak, corresponding to an annihilation process, are only in the lower part,
- plot two spectra:
	- o number of counts vs. channel, corresponding to energy (if a reliable energy calibration is available),
	- o number of counts vs. time for each energy range as well as for a full energy range,
- all results of the analysis are saved in text files.





### **3.1. Description of the AU\_main program used for data analysis**

The AU main program to analyze data from Gamma Camera (GC) is written in  $C_{++}$ language. It consists of classes:

*Events*

*ReadBin*

*PileupFilter*

*Analysis*

The DAQ time to start acquisition is provided by the JET absolute time (CTTS clock and trigger) with the precision of 1 MHz. The data file is composed of events (PWIDTH samples of 16-bit). Each event corresponds to: an occurrence time (64-bit) which is the first 4 samples and a pulse amplitude which is (PWIDTH-4) samples.

There are three structures implemented due to three types of saved data: *struct event\_CeBr\_1024{ uint64\_t time; int16\_t amplitude[1024]; }; struct event\_CeBr\_512{ uint64\_t time; int16\_t amplitude[512];}; struct event\_CsI{ int32\_t amplitude; uint32\_t time; };*

The class *Events* is a template of a vector which contains such information about signals as a time and an amplitude. This object is created by the function *readEventsFromBin()* which is a part of the class *ReadBin*.

The class *ReadBin* is used to read a binary output file which is generated by DAQ dedicated for the GC. This class is able to distinguish three types of events:

- 1. events collected with CsI scintillatrs,
- 2. events collected with a new CeBr<sub>3</sub> scintillator and an old DAO
- 3. events with a new crystal LaBr3:Ce and a new DAQ. In this case, the *ReadBin* class creates the *PileupFilter* object to unfold pile-up events. The function *readEventsFromBin()* returns a vector of *Event*s objects which is sent to the object *Analysis*.

The class *PileupFilter* is used to unfold pile-up signals. Two types of pile-up events are considered:

- a signal length is longer than the data acquisition time (including the detector, MPPC and electronics resolution time),
- a signal rate is relatively high.

In each of these two signal types, more than one signal is registered in an event, see Fig. 2.







Fig. 2. Signals registered in the GC detectors: left – single event, right – pile-up event.

In case of a single event, a signal shape is described by a formula:

$$
f(t) = a \cdot (\exp(-t/\tau_1) - \exp(-t/\tau_2)),
$$
 Eq. (1)

where *a* is a constant, *t* is time,  $\tau_1$  and  $\tau_2$  are time constants dependent on the detector. A time length and an amplitude of this signal can be easily calculated, see Fig. 3. In addition, it is possible to obtain a pattern of such a single signal.



Fig. 3. Fit, described by Eq. (1), to a measured signal: a= -1333,  $\tau_1$ =16.99,  $\tau_2$ =84.26.

In case of a pile-up event, more than one consists of more than signal is registered. So, a more complicated special formula is needed to obtain both time and amplitude of a signal. This is realized by the class *PileupFilter*.





When a pile-up event is registered, it is compared to the single signal pattern and a pile-up event is decomposed into single ones. Such extracted new single signals are saved. For each event, a pile-up discriminator value, equal to a ratio of an area below a signal to a rectangular frame area based on a signal amplitude and a number of samples, was calculated, see blue rectangle in Fig. 2.

The *Analysis* object loads the *Event* object and builds an energy spectrum and calculates how many events were recorded during a shot, especially for energy  $\leq 520$  keV and  $> 520$  keV.



The flowchart of the analysis program is shown in Fig. 3.

Fig. 3. Flowchart of the analysis program.

Three input parameters should be supplied by the user:

- 1. shot number,
- 2. detector channel number from 1 to 19: 1-10 for the horizontal camera and 11-19 for the vertical one,
- 3. time step in seconds as demanded by the user.

Staring the AU\_main program: *./AU\_main 92449 10 0.001*





A *ReadBin* object calls *readFromFile()* which uses a *PileupFilter* object and returns an *Event* object which collects all events. The *Event* object, with all collected events, is then relayed to an *Analysis* object, calling dedicated analysis methods.

The output of the AU\_main analysis program contains the following four text files of the same structure: channel, energy in keV, counts

- 1. Analysis\_RE\_Spectrum.dat
- 2. Analysis\_RE\_low.dat for the low energy range  $\leq 520$  keV
- 3. Analysis\_RE\_high.dat for high energy range >520 keV
- 4. Analysis RE full.dat for a full energy range

The AU\_main program is prepared for a Linux operating system and extensively tested on JAC computers.

### **3.2. Comparison of CsI and CeBr<sup>3</sup> detector performance**

CsI scintillators have slower response time and lower efficiency in comparison to CeBr3 scintillators. In Fig. 4 a distribution of counts vs. time from a shot #91069 is presented with a time step equal to 0.001 second (input to the AU\_main program).





No signal with old CsI is registered during a disruption. This effect can be related to a higher dead time of the CsI detector which was not able to register higher count rates. This effect





was observed in each channel with old CsI-based detectors, i.e. from channel #1 to #8 in a horizontal part and from #11 to #19 in a vertical part of the Gamma-ray Camera.

## **3.3. Analysis of shots from 2016 campaign at JET**

The CeBr3-based detector was installed at JET in May 2015 in the channel #10 in the Gamma-ray Camera.

The following parameters were defined for shots, see Fig. 5, :

- 1. *Tstart* time when gamma events start to be registered in the detector,
- 2. *Tend* time when no more gamma events end to be registered in the detector,
- 3. *ΔT=Tend-Tstart* the duration of a disruption,
- 4. *N\_full* –number of counts during *ΔT* in the full energy range,
- 5. *N\_low* –number of counts during *ΔT* in the low energy range,
- 6. *N\_high*–number of counts during *ΔT* in the high energy range.
- 7. *N\_full/* $\Delta T$  *–* average number of counts during  $\Delta T$  in the full energy range,
- 8. *N\_low/ΔT* average number of counts during *ΔT* in the low energy range,
- 9. *N\_high/ΔT* average number of counts during *ΔT* in the high energy range.



Fig. 5. Definition of parameters defined for each shot.

Results of analysis performed with AU\_main program are presented in Table 1.

Below Table 1, distributions plotted for each shot show:





- 1. amplitude vs. DAQ channel (upper),
- 2. time distribution of events in the full, low and high energy range (lower).

































**#91069**. 1.5MA / 3.0T. DMV1 argon 7.6 bar @60.0s. DMV2 Krypton 45 bar @60.039s. Good! Runaway plateau ends at 60.065s. The end is not exactly in the middle between the 2 previous pulses, but it still goes in the right direction. 700 degC impact on the dump plate from KL7































 **#91079**. Repeat #91069 to estimate the variability of the runaway beam. 1.5MA / 3.0T. DMV1 argon 7.6 bar @60.0s. DMV2 Krypton 45 bar @60.039s. Longer than #91069, but starts with a higher initial current. Still, it is longer than #91071 which was not mitigated. Annoying. The scan is now all over the place, and this runaway beam just seems to do what it wants.







 **#91081**. Repeat #91068 with Neon max pressure. 1.5MA / 3.0T. DMV1 argon 7.6 bar @60.0s. DMV2 Neon 45 bar @60.028s. Beam ends around 60.065s. Longer than #91068, but not enough to conclude that Neon is less efficient than krypton. And another unconclusive pulse.















































































#### **The report was prepared by the NCBJ team**

Arkadiusz Urban, Andrzej Broslawski and Izabella Zychor.

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